expansions for the forces, moments, and transverse forces. Let $\varepsilon_{\theta}$ denote the circumferential strain. If $\varepsilon_{\theta}\left(W_{n}\right) \neq 0, n=0,1$, then the asymptotic expansions of the moment $M_{\theta}$ and the transverse forces $N_{r}$, $N_{\theta}$ start with the power $\varepsilon^{-2}$.

For $\varepsilon_{\theta}\left(\mathrm{w}_{\mathrm{n}}\right)=0, \mathrm{n}=0,1$ the plate is inextensible in the circumferential direction.
In conclusion, we note that the asymptotic for the problem of bending a symmetricaliy assembled anisotropic rectangular laminar she11 [5] under strictly nonzero steepness of the family of bonding fibers can be constructed completely analogously (in a formal complication of the computations).

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STABILITY OF MAGNETIC SUSPENSION IN A DIRECT-CURRENT MAGNETIC FIELD
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The problem of suspension of a body for a lengthy period of time using permanent magnets has attracted the interest of researchers. A detailed bibliography of studies of this problem, an analysis of the state of the art, and original results have been presented in [1, 2].

The major result achieved has been Ernshaw's theorem, which indicates the instability of such suspension. However this theorem is concerned with steady state situations, and as we will demonstrate below, is inapplicable to dynamic systems.

1. We will consider the configuration of magnets shown in Fig. 1. We will consider motion of an infinitely long rod in the magnetic channel along the axis $0 x$. The weight of the $\operatorname{rod} \mathrm{P}=\mathrm{mg}$ is compensated by magnets of one sign 1 or 2 . Along the channel sides there is a system of permanent magnets of alternating polarity, which interacts with a similar system located on the rod. We will assume that the pole step along the axis $0 x$ is equal to $\lambda=2 \pi / \mathrm{k}$, where $k$ is the wave number. We will assume the magnetic material to be saturated with a value of $\mu=1$ (where $\mu$ is the relative permittivity), as in a vacuum. Considering further that magnet system 3 has a vertical length, we will neglect forces produced by interaction of magnets 3 and 4 during vertical oscillations of the rod.

We will now perform some preliminary calculations. At the point $M_{0}\left(x_{0}, y_{0}, z_{0}\right)$ let there be some magnetic charge $q$. Its potential at the point $M(x, y, z)$ is equal to $U=q / 4 \pi \mu_{\circ} r$, $r^{2}=\left(x-x_{0}\right)^{2}+\left(y-y_{0}\right)^{2}+\left(z-z_{0}\right)^{2}$, where $\mu_{0}$ is the absolute magnetic permittivity of free space. The force produced by interaction of two charges $q^{+}, q^{-}$located at these points is given by the expression

$$
F=\frac{1}{4 \pi \mu_{0}} \frac{q^{+} q-}{r^{-}}
$$

and is directed along the vector joining the charges.
We will consider an infinite magnetic pole located along the axis $O z$. For an element of the pole $d z$ the magnetic charge is equal to $d q=\gamma^{+} d z$, where $\gamma^{+}$is the linear charge density. The force of interaction with an analogous elementary charge sectioned from another magnetic

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pole will be:

$$
d^{2} F=\gamma^{+} \gamma^{-} d z d z_{0} /\left(4 \pi \mu_{\mathbf{0}} r^{2}\right)
$$

We will now define the force with which a charge located in the plane $y=\delta$ is attracted to a pole the edge of which lies in the plane $y=-\delta$ at coordinate $x$. Projecting the force $d^{2} F$ onto the shortest line joining the point $M_{0}\left(x_{0}, y_{0}=\delta, z_{0}\right)$ to the pole $x^{\prime}=x, y^{\prime}=-\delta, z^{\prime}=$ $z$, we have

$$
d^{2} R=d^{2} F \frac{\sqrt{\left(x-x_{0}\right)^{2}+4 \delta^{2}}}{\sqrt{\left(x-x_{0}\right)^{2}+4 \delta^{2}+\left(z-z_{0}\right)^{2}}}=\frac{\gamma^{4} \gamma^{-}}{4 \pi \mu_{0}} \frac{\sqrt{\left(x-x_{0}\right)^{2}+4 \delta^{2}}}{\left\{\left(x-x_{0}\right)^{2}+4 \delta^{2}+\left(z-z_{0}\right)^{2}\right\}_{0}^{3 / 2}} .
$$

The force with which an elementary charge is attracted by an infinite pole is given by the integral

$$
\begin{equation*}
d R=\frac{\gamma^{+} \gamma^{-} d z}{4 \pi \mu_{0}} \int_{-\infty}^{\infty} \frac{\left\{\left(x-x_{0}\right)^{2}+4 \delta^{2}\right\} d z}{\left\{\left(x-x_{0}\right)^{2}+4 \delta^{2}+\left(z-z_{0}\right)^{2}\right\}^{3 / 2}} \tag{1.1}
\end{equation*}
$$

Integrating Eq. (1.1) and performing the replacement $z-z_{0}=u \sqrt{\left(x-x_{0}\right)^{2}}+4 \delta^{2}$, we obtain [3]

$$
d R=\gamma^{+} \gamma^{-} d z_{0} /\left[2 \pi \mu_{0} \sqrt{\left.\left(x-x_{0}\right)^{2}+4 \delta^{2}\right]}\right.
$$

A pole of unit length along the axis $0 z$ will experience an attractive force

$$
\begin{equation*}
R=\frac{\gamma^{+} \gamma^{-}}{2 \pi \mu_{0}} \frac{1}{\sqrt{\left(x-x_{0}\right)^{2}+4 \delta^{2}}} \tag{1.2}
\end{equation*}
$$

In_the case where the charge is distributed along the axis $0 x$, i.e., $d q=q^{-} d x$ (where $q^{-}=$ $\mathrm{d} \gamma^{-} / \mathrm{dx}$ is the magnetic charge density) we write $\mathrm{Eq} .(1.2)$ in the form

$$
d^{2} R=q^{+} q^{-} d x d x_{0} /\left[2 \pi \mu_{0} \sqrt{\left.\left(x^{\prime}-x_{0}\right)^{2}+4 \delta^{2}\right]} .\right.
$$

The projections of this force on the $O x$ and $O y$ axes are equal to

$$
\begin{align*}
d^{2} X & =\frac{\left(x-x_{0}\right) q^{+} q^{-} d x d x_{0}}{2 \pi \mu_{0}\left\{\left(x-x_{0}\right)^{2}+4 \delta^{2}\right\}}  \tag{1.3}\\
d^{2} Y & =\frac{2 \delta q^{+} q^{-} d x d x_{0}}{2 \pi \mu_{0}\left\{\left(x-x_{0}\right)^{2}+4 \delta^{2}\right\}} .
\end{align*}
$$

First of all, we integrate Eq. (1.3) over the variable $x$. By doing this we define the force produced upon an elementary magnetic charge by all charges in the plane $y=-\delta$ :

$$
\begin{align*}
& d X=\frac{q^{+}\left(x_{0}\right) d x_{0}}{2 \pi \mu_{0}} \int_{-\infty}^{\infty} \frac{\left(x-x_{0}\right)^{2} q^{-}(x) d x}{\left(x-x_{0}\right)^{2}+4 \delta^{2}}, \\
& d Y=\frac{q^{+}\left(x_{0}\right) d x_{0}}{2 \pi \mu_{0}} \int_{-\infty}^{\infty} \frac{2 \delta q^{-}(x) d x}{\left(x-x_{0}\right)^{2}+4 \delta^{2}} \tag{1.4}
\end{align*}
$$

Performing the substitution $x-x_{0}=u$ in Eq. (1.4) and taking the magnetic charge density in the form

$$
q^{+}\left(x_{0}\right)=q_{0} \sin k\left(x_{0}+\xi\right), q^{-}(x)=q_{0} \sin k x
$$

we find, by symmetry, that

$$
\begin{aligned}
& d X=\frac{q_{0}^{2}}{\pi \mu_{0}} \sin k\left(x_{0}+\xi\right) \cos k x_{0} d x_{0} \int_{0}^{\infty} \frac{u \sin k u d u}{u^{2}+4 \delta^{2}} \\
& d Y=\frac{q_{0}^{2}}{\pi \mu_{0}} \sin k\left(x_{0}+\xi\right) \sin k x_{0} d x_{0} \int_{0}^{\infty} \frac{2 \delta \cos k u d u}{u^{2}+4 \delta^{2}}
\end{aligned}
$$

Since according to [3]

$$
\int_{0}^{\infty} \frac{2 \delta \cos k u d u}{u^{2}+4 \delta^{2}}=\int_{0}^{\infty} \frac{u \sin k u d u}{u^{2}+4 \delta^{2}}=\frac{\pi}{2} t^{-2 \delta k}
$$

we have

$$
d X=\frac{q_{0}^{2} \mathrm{e}^{-2 \delta k}}{2 \mu_{0}} \sin k\left(x_{0}+\xi\right) \cos k x_{0}^{d x}, \quad d Y=\frac{q_{0}^{2} \mathrm{e}^{-2 \delta k}}{2 \mu_{0}} \sin k\left(x_{0}+\xi\right) \sin k x_{0} d x_{0}
$$

From this we define the attractive force per unit surface:

$$
X=\frac{q_{0}^{2} e^{-2 \delta \hbar}}{2 \mu_{0}} \lim \frac{k}{\pi n} \int_{0}^{\frac{\pi n}{k}} \sin k\left(x_{0}+\xi\right) \cos k x_{0} d x_{0}=\frac{q_{0}^{2} \mathrm{e}^{-2 \delta k}}{2 \mu_{0}} \sin / \hbar \xi \lim \frac{k}{\pi n} \int_{0}^{\frac{\pi n}{h}} \cos ^{2} k x_{0} d x_{0}
$$

i.e.,

$$
\begin{equation*}
X=\frac{q_{0}^{2} e^{-2 \delta k}}{2 \mu_{0}} \sin k \xi, \quad Y=\frac{q_{0}^{2} e^{-2 \delta k}}{2 \mu_{0}} \cos k \tag{1.5}
\end{equation*}
$$

We will now consider the force of attraction (or repulsion) between the surface $y=\delta$ and the surface $y=-\delta-\eta$, formed by the two ends of the magnets. This force can be calculated from Eq. (1.5) by replacing the quantity $2 \delta$ by $2 \delta+Z$ and considering the sign of the force. Then the total force is given by the expression

$$
X=\frac{q_{0}^{2} \mathrm{e}^{-2 \delta k}}{2 \mu_{0}} \cdot\left(1-\mathrm{e}^{-i k}\right)^{2} \sin k \xi, \quad Y=\frac{\dot{q}_{0}^{2} \mathrm{e}^{-3 \delta k}}{2 \mu_{0}}\left(1-\mathrm{e}^{-i k}\right)^{2} \cos k_{\xi}
$$

Considering that the force of interaction between a unit magnetic pole dq and a magnetic fiel: of intensity $H$ is determined by the product $d X=$ dqH, from Eg. (1.5) we find the value of the field intensity on the surface itself, at the point corresponding to the maximur

$$
H_{x}=q_{0} / \mu_{0}, \quad H_{y}=q_{0} / \mu_{0}
$$

Using the magnetic induction $B=\mu_{0} H$, we write the force

$$
\begin{equation*}
X=\frac{E^{2}}{2 \mu_{0}} \mathrm{e}^{-26 k}\left(1-\mathrm{e}^{-l k}\right)^{2} \sin I \xi, \quad Y=\frac{B^{2}}{2 \mu_{0}} e^{-2 \delta k}\left(1-e^{-l k}\right)^{2} \cos k \xi \tag{1.6}
\end{equation*}
$$

2. Knowing the force characteristics of the magnetic field of system (1.6), we can construct dynamic equations. We will displace the rod to the right by an amount $y$. Then the right-hand gap will be $\delta-y$, and the left-hand, $\delta+y$. In this case the attractive (repulsive) force for the right-hand gap will be

$$
Y^{+}=-\frac{B^{2}}{2 \mu_{0}} \mathrm{e}^{-2 \delta h-y k}\left(1-\mathrm{e}^{-l k}\right)^{2} \cos k x
$$

while for the left-hand gap

$$
Y^{-}=\frac{B^{2}}{2 \mu_{0}} \mathrm{e}^{-2 \delta k+\vartheta k}\left(1-\mathrm{e}^{-l h}\right)^{2} \cos k x
$$

Then the total force will be given by

$$
\mathscr{G}=Y^{+}+Y^{-}=\frac{B^{\AA}}{\mu_{0}} \mathrm{e}^{-2 \delta k}\left(1-\mathrm{e}^{-l k}\right)^{2} \cos k x \operatorname{sh} k y
$$

Similarly we define the force $\mathscr{P}$ by the expression

$$
\begin{equation*}
\mathscr{X}=X^{+}+X^{-}=\frac{B^{2}}{\mu_{0}} \mathrm{e}^{-2 \delta k}\left(1-\mathrm{e}^{-l h}\right)^{2} \sin k x \operatorname{ch} k y \tag{2.1}
\end{equation*}
$$

Aside from the force $\mathscr{G}$, there will act a destabilizing force $\mathscr{G}^{*}$, produced by the lower magnetic system, the supporting mechanism. For the case where this support system is composed of several magnetic poles, by analogy to Eq. (2.1) we find for the force $\mathscr{G}^{*}$

$$
\mathscr{G} *=\frac{B_{1}^{2}}{\mu_{0}} \mathrm{e}^{-2 \delta k_{1}} \sin k_{1} y
$$

Here $k_{1}$ is the wave number of the lower magnetic system. The supporting force will be equal to

$$
P=\frac{B_{1}^{2}}{\mu_{0}} e^{-2 \delta k_{1}} \cos k_{1} y
$$

At $y=0$ we have $P=m g$. Hence we find $m g=\left[\exp \left(-2 \delta k_{1}\right)\right] B_{1}^{2} / \mu_{0}$, then $\mathscr{G}^{*}=m g$ sin $k_{1} y$. With consideration of all forces acting we can write the equations of motion of the rod:

$$
\begin{gather*}
m \frac{d^{2} x}{d t^{2}}=\frac{B^{2}}{\mu_{0}} \mathrm{e}^{-2 \delta k}\left(1-\mathrm{e}^{-l k}\right)^{2} \sin k x \operatorname{ch} k y,  \tag{2.2}\\
m \frac{d^{2} y}{d t^{2}}=-\frac{B^{2}}{\mu_{0}} \mathrm{e}^{-2 \delta k}\left(1-\mathrm{e}^{-l k}\right)^{2} \cos k x \operatorname{sh} k y+m g \sin k_{1} y .
\end{gather*}
$$

Introducing the notation

$$
\frac{B^{2}}{\mu_{0} m} \mathrm{e}^{-2 \delta k}\left(1-\mathrm{e}^{-l k}\right)^{2}=A, \quad k_{1}=i
$$

we transform Eq. (2.2) to the form

$$
\begin{equation*}
d^{2} x / d t^{2}=A \sin k x \operatorname{ch} k y, d^{2} y / d t^{2}=-A \cos k x \operatorname{sh} k y+g \sin l y \tag{2.3}
\end{equation*}
$$

It can be shown that Eq. (2.3) is the equation of motion

$$
d^{2} x_{1} / d t^{2}=-\partial \Pi(x) / \partial x_{1}, d^{2} x_{2} / d t^{2}=-\partial \Pi(x) / \partial x_{2}
$$

in a force field with potential $\Pi(x)=a \operatorname{ch} k x_{2} \cos k x_{1}+b \cos \tau_{x_{2}}\left(x_{1}=x, x_{2}=y\right)$ and permits definition of a Hamiltonian energy integral.

$$
\begin{equation*}
H=\frac{1}{2}\left(y_{1}^{2}+y_{2}^{2}\right)+a \operatorname{ch} k x_{2} \cos k x_{1}+b \cos l x_{2} \tag{2.4}
\end{equation*}
$$

where $y_{1}, y_{2}$ are the canonical momenta corresponding to the coordinates $x_{1}, x_{2}$.
We will determine the singular points of system (2.3) (equilibrium positions) and estab1ish conditions for their stability by studying the topological structure of the energy surface defined by Hamiltonian (2.4) in their vicinity, using the approach described in [4, 5]. We will consider the equations


Fig. 2

$$
\begin{equation*}
a k \operatorname{ch} k x_{2} \sin k x_{1}=0,-a k \operatorname{sh} k x_{2} \cos k x_{1}+b l \sin l x_{2}=0 \tag{2.5}
\end{equation*}
$$

From the first equation we obtain the solutions $x_{1}=\pi i / k(i=0, \pm 1, \pm 2, \ldots)$, with consideration of which the second equation of Eq . (2.5) takes on the form

$$
\begin{equation*}
-(-1)^{i} a k \operatorname{sh} k x_{2}+b l \sin l x_{2}=0 \tag{2.6}
\end{equation*}
$$

We will study the solutions of Eq. (2.6) graphically. Thus, for $x_{1}=\pi i_{e} / k\left(i_{e}=0, \pm 2\right.$, ...) we write Eq. (2.6) in the form $b l \sin ~ Z x_{2}=a k \operatorname{sh} k x_{2}$. Graphing the functions $f_{1}=$ $b l \sin Z x_{2}, E_{2}=a k \operatorname{sh} \mathrm{kx}_{2}$, at the points of their intersection we obtain solutions of Eq. (2.6). It is evident from Fig. 2 that solutions other than $x_{2}=0$ will exist only when the inequality

$$
\left.\left(b l \sin l x_{2}\right)^{\prime}\right|_{x_{2}=0}>\left.\left(a k \operatorname{sh} k x_{2}\right)^{\prime}\right|_{x_{2}=0}
$$

is satisfied. The total number of solutions of Eq. (2.6), $N$, can be determined in the following manner. Let $n$ be the largest number of waves for which the inequality

$$
b l>a k \operatorname{sh} k(\pi n / l+\pi / 4 l)
$$

is satisfied. Hence

$$
\begin{equation*}
n=[-1 / 4+(l / \pi k) \operatorname{arcsh} \sigma], \quad \sigma=b l / a i \tag{2.7}
\end{equation*}
$$

Then $N=(2 n+2) 2-1=4 n+3$. For $x_{1}=\pi i_{o} / k\left(i_{0}=+1, \pm 3, \ldots\right)$ the number $N_{1}$ of solutions of Eq. (2.6), i.e., equilibrium positions, is the same, or two smaller, i.e., $N_{1}=N$ or $N_{1}=$ $N-2$. Thus we have the following discrete set of equilibrium positions:

$$
\begin{aligned}
& G_{0 e}=\left\{x, y ; y_{1}=y_{2}=0 ; x_{1}=\pi i_{e} / k, x_{2}=0 ; i_{e}=0, \pm 2, \ldots\right\} \\
& G_{00}=\left\{x, y ; y_{1}=y_{2}=0 ; x_{1}=\pi i_{0} / k, x_{2}=0 ; i_{0}= \pm 1, \pm 3, \ldots\right\} \\
& G_{j_{i}}=\left\{x, y ; y_{1}=y_{2}=0 ; x_{1}=\pi i / k ; x_{2}=x_{2 j}(\gamma, \sigma) ; i=0, \pm 1, \pm 2, \ldots ; j= \pm 1,\right. \\
& \pm 2 \ldots\} .
\end{aligned}
$$

We note the significance of the parameters $\sigma$ and $\gamma$ in existence of a sequence of equilibrium positions for Eq. (2.3).

We will determine the type of topological structure of the energy surface $H=H(x, y)$ by examining the Gaussian curvature defined by the expression
where

$$
\begin{gather*}
G(x, y)=\left(1+|\operatorname{grad} H|^{2}\right)^{-3 / 2} \operatorname{det} \operatorname{Hess}(I F ; x, y), \\
\text { Hess }(H ; s)=\left(\frac{\partial^{2} I I(s)}{\partial s \partial s}\right) \tag{2,8}
\end{gather*}
$$

( $s$ is the variable system vector, $s=\left(x_{1}, x_{2}, y_{1}, y_{2}\right)$ ). We will use matrix (2.8) to study the Gaussian curvature.

We will now make use of the following
Definition. The surface $H(x, y)$ is topologically equivalent to the surface $H^{*}(x, y)$ in some related region $\Gamma(x, y)$ if the following relationship is satisfied:

$$
\begin{equation*}
Q^{\mathrm{T}}(x, y) \text { Hess } \quad(H ; x, y) Q(x, y)=\operatorname{Hess}\left(H^{*} ; x, y\right), \tag{2.9}
\end{equation*}
$$

where $Q(x, y)$ is some quadratic matrix.
In the case where the region is the vicinity of an equilibrium position, each of which in the given case is a nondegenerate critical point of the surface $H(x, y)$, and the surface $H^{*}(x, y)$ is canonical, i.e., $H^{*}(x, y)=\sum_{i=1}^{n}\left(\alpha_{i} z_{i}^{2}+\beta_{i} y_{i}^{2}\right)$, where $\alpha_{i}, \beta_{i}$ take on the values $\pm 1$, the given definition is based on Morse's lemma [6].

Since the matrix Hess ( $H$; $x, y$ ) is symmetric, a transform (the matrix $Q(x, y)$ ) will always exist. However if $H \%(x, y)$ is such that the matrix Hess ( $H$; $x, y$ ) is diagonal while the region $\Gamma(x, y)$ is not local, then the question of definition of the transform $Q(x, y)$ remains open.

It is known [4] that if the condition $\Delta_{r}(x, y)>0$ for all $i=\overline{1,2 n}$, where $\left\{\Delta_{r}\right\}$ is the sequence of negative main minors of matrix (2.8), is satisfied, then $G(x, y)$ is positively defined. It is then possible to determine a sequence (with consideration of the existence of a sequence of equilibrium positions for Eq. (2.3)) of closed compact invariant regions

$$
\begin{aligned}
& \bar{\Omega}_{i \alpha}^{+}(x, y)=\Omega_{i \alpha}^{+} \cup \partial \Omega_{i \alpha}^{+}, \partial \Omega_{i \alpha}^{+}-\text {is the boundary of the region, } \\
& \Omega_{i \alpha}^{+}=\Omega_{i}^{+} \cap\left\{H(x, y)<\alpha=H_{\mathrm{inf}}^{\partial \Omega_{i}^{+}}=\inf _{\partial \Omega_{i}^{+}} H\right\}, \\
& \Omega_{i}^{+}=\left\{\begin{array}{l}
2 n \\
\bigcap_{r=1}^{+} \\
\Omega_{i r}
\end{array}\right\}, \quad \Omega_{i r}^{+}=\left\{\Delta_{i r}>0\right\}, i-\text { is the number of the region, }
\end{aligned}
$$

containing both a stable equilibrium position and periodic (quasiperiodic) system trajectories. In each such region Hamiltonian (2.4) is topologically equivalent to the Hamiltonian

$$
I^{*}=\frac{1}{2}\left(y_{1}^{2}+y_{2}^{2}\right)+\frac{1}{2}\left(x_{1}^{2}+x_{2}^{2}\right) .
$$

We will now consider the cases of indeterminate sign and negative sign determination of $G(x, y)$ (and thus matrix (2.8)).

For the given Hamiltonian (2.4) the matrix (2.8) has the simple form

$$
\operatorname{Hess}(H ; x)=\left(\begin{array}{cc}
\left(\frac{\partial^{2} \Pi(x)}{\partial x_{i} \partial x_{j}}\right. & (0)  \tag{2.10}\\
(0) & (1)
\end{array}\right),
$$

where its elements are $2 \times 2$-matrices. Because of the form of matrix (2.10) it is sufficient to examine its submatrix

$$
\left(\frac{\partial^{2} \Pi(x)}{\partial x_{i} \partial x_{j}}\right)=\left(\begin{array}{ll}
-a k^{2} \cos k x_{1} \operatorname{ch} k x_{2} & -a k^{2} \sin k x_{1} \operatorname{sh} k x_{2}  \tag{2.11}\\
-a k^{2} \sin k x_{1} \operatorname{sh} k x_{2} & a k^{2} \cos k x_{1} \operatorname{ch} k x_{2}-b l^{2} \cos l x_{2}
\end{array}\right),
$$

which corresponds to study of the topology of the surface $\Pi=\Pi(x)$. Using Eq. (2.11), we define

$$
\begin{aligned}
& \Delta_{1}(x)=-a k^{2} \cos k x_{1} \operatorname{ch} k x_{2}, \\
& \Delta_{2}(x)=a^{2} k^{4}\left(-\operatorname{ch}^{2} k x_{2}+\sin ^{2} k x_{1}+\gamma \cos k x_{1} \cos l x_{2} \operatorname{ch} k x_{2}\right),
\end{aligned}
$$

while

$$
\frac{\Delta_{2}(x)}{\left(1+\left(\frac{\partial \Pi(x)}{\partial x_{1}}\right)^{2}+\left(\frac{\partial \Pi(x)}{\partial x_{2}}\right)^{2}\right)^{3 / 2}}=G(x),
$$

where $G(x)$ is the Gaussian curvature of the surface $\Pi=\Pi(x)$.
For points of the sets $G_{o e}$ and $G_{o o}$ we have

$$
\begin{aligned}
& \Delta_{1}\left(M_{0 e}\right)=-a k^{2}, \quad \Delta_{2}\left(M_{0 e}\right)=a^{2} k^{4}(-1+\gamma) ; \\
& \Delta_{2}\left(M_{00}\right)=a k^{2}, \quad \Delta_{2}\left(M_{00}\right)=a^{2} k^{4}(-1-\gamma) .
\end{aligned}
$$

From this it follows that for $\gamma<1$ all equilibrium positions of the sets $G_{o e}$, $G_{o o}$ considered in the subspace $X$ are hyperbolic, and the structure of the surface defined by the potential energy function in the vicinity of these points is equivalent to a saddle. Therefore Hamiltonian (2.4) in the vicinity of the equilibrium position of the sequences Goe and Goo is topologically equivalent to Hamiltonian

$$
\begin{equation*}
H^{*}=-\frac{1}{2}\left(y_{1}^{2}+y_{2}^{2}\right)+\frac{1}{2}\left(\mp x_{1}^{2} \pm x_{2}^{2}\right) \tag{2.13}
\end{equation*}
$$

where the upper sign refers to points of the set $G_{o e}$, and the lower sign to Goo. At $\gamma>1$ at the equilibrium points of Goe matrix (2.11) is negatively determined. From this it follows that at these points the function $\Pi(x)$ has a maximum, so that Hamiltonian (2.4) in the vicinity of each of these points is topologically equivalent to the Hamiltonian

$$
\begin{equation*}
H^{*}=\frac{1}{2}\left(y_{1}^{2}+y_{2}^{2}\right)-\frac{1}{2}\left(x_{1}^{2}+x_{2}^{2}\right) \tag{2.14}
\end{equation*}
$$

For all remaining (at $\gamma>1$ ) equilibrium positions we obtain

$$
\Delta_{1}\left(M_{j i}\right)=-(-1)^{i} a k^{2} \operatorname{ch} k x_{2}, \Delta_{2}\left(M_{j i}\right)=a^{2} k^{4} \operatorname{ch} k x_{2}\left(-\operatorname{ch} k x_{2}+(-1)^{i} \gamma \cos l x_{2}\right)
$$

Hence at $i=i_{0}$ it follows that $\Delta_{1}\left(M_{j i_{0}}\right)>0, \quad \Delta_{2}\left(M_{j i_{0}}\right)=a^{2} k^{4} \operatorname{ch} k x_{2}\left(-\operatorname{ch} k x_{2}-\gamma \cos l x_{2}\right)$, and for sufficiently large $\gamma$ at $j=j o$ satisfaction of the inequality $\Delta_{2}\left(M_{j_{o} i_{0}}\right)>0$ is possible. Therefore, aside from maximum points and saddle points, in the vicinity of which Eqs. (2.14), (2.13) are valid, at $i=i_{0}, j=j o$ there exist points where the function $\Pi$ (s) is at a minimum, thus allowing representation of Hamiitonian (2.4) in their vicinity in the form of Eq. (2.9) so that these are stable equilibrium points of system (2.3).

Thus, the model admits the existence of magnetic traps in a manner similar to the way the limited model three-body problem of celestial mechanics admits existence of gravitational traps corresponding to the libration points $L_{4}$ and $L_{5}$ [7]. Here the existence of stable equilibrium positions contradicts Ernshaw's theorem [1] which states the impossibility of a stable configuration created by permanent magnets, and can be explained by the lack of consideration of displacement along the $z$ coordinate, i.e., by the limited formulation of the problem.

Our goal now will be to establish conditions for existence of trajectories corresponding to longitudinal motion (along the axis $0 x_{1}$ ) with limited change in the coordinates $y_{1}, x_{2}, y_{2}$ and to study the stability of such a trajectory with respect to small initial perturbations. We can relate the search for such conditions to establishing the possibility of representing Hamiltonian (2.4) in some tube $T(x, y)=\left\{x, y:\left|x_{1}\right|<\infty,\left|x_{2}\right|<\Delta_{0}, \Delta_{1}<H(x, y)<\Delta_{2}\right.$, $\Delta_{1}<$ $C=$ const $\}$ in the equivalent form $H *(x, y)$, the trajectories of which have the required properties, and then, to establishment of time limits or other conditions for maintenance of these properties.

Consideration of the trajectories of Hamiltonian (2.4) at $\gamma<1$ on the basis of an equivalent $H^{*}(x, y)$ of the form of $E q$. (2.13) in the tube $T$ shows the impossibility of existence in system (2.3) of trajectories corresponding to the limiting motion. Therefore we will turn to consideration of the case $\gamma>1$. Here the topological structure of the force field (potential energy) is insured by high parameter values ( $\gamma=62.5, \sigma=25.0$ ). By Eq. (2.7) the quantities $\gamma$ and $\sigma$ define the value $n=1$. Below we will show that the condition $\gamma \gg 1$, $n \geqslant 1$ is significant in determining the possibility of reduction to equivalent structures.

By using reductions of the relative character of equilibrium positions it can be established that upon change in the coordinates $x_{1}, y_{1}, x_{2}, y_{2}$ in the tube

$$
T(x, y)=\left\{x, y:\left|x_{1}\right|<\infty, \quad\left|x_{2}\right|<x_{2}\left(M_{1 i_{e}}\right), \quad \Pi\left(M_{1 i_{e}}\right)<\Pi(x, y)<\Pi\left(M_{0 i_{e}}\right)\right\}
$$

where $M_{1} i_{e}$ is the second point in the sequence of equilibrium positions, corresponding to $i_{e}$, Moie is the first point of the same sequence, the maximum point of the function $\Pi$ ( $x$ ),

Hamiltonian (2.4) is topologically equivalent to Hamiltonian

$$
\begin{equation*}
I^{*}=\frac{1}{2}\left(y_{1}^{2}+y_{2}^{2}\right)+\frac{1}{2}\left(\cos x_{1}+\cos x_{2}\right) \tag{2.15}
\end{equation*}
$$

Hamiltonian (2.15) can be analyzed easily. From Eq. (2.15) we obtain two independent systems $\mathrm{H}_{1}$ and $\mathrm{H}_{2}$, where

$$
H_{1}=\frac{1}{2}\left(y_{1}^{2}+\cos x_{1}\right), \quad H_{2}=\frac{1}{2}\left(y_{2}^{2}+\cos x_{2}\right)
$$

Hence, if the values of the energy $H_{1}$ are such that the initial point ( $x_{1}^{0}, y_{1}^{0}$ ) is located beyond the separatrix dividing the region of periodic motions of the system $H_{1}$ from the region of motions departing to infinity along the axis $O x_{1}$ at finite $y_{1}$, and the values of energy $H_{2}$ are such that the initial point ( $\mathrm{x}_{2}^{0}, y_{2}^{0}$ ) lies within the region of periodic motions of the system $H_{2}\left(H_{2}<H_{2} C\right.$, where $H_{2} C$ is the energy corresponding to motion along the separatrix), the trajectory obtained from the initial point ( $x_{1}^{0}, y_{1}^{0}, x_{2}^{0}, y_{2}^{0}$ ) will correspond to the desired one. Naturally, since the given system (2.3) is bound (which leads to redistribution of the initial energy between subsystems $H_{1}$ and $H_{2}$ while maintaining the integral $H=H_{1}+H_{2}$ ), there must exist a moment of time $t^{*}$ at which the energy $\mathrm{H}_{2} \mathrm{C}$ in system $\mathrm{H}_{2}$ will be exceeded. Nevertheless, it can be expected that the interval [0, t*] will not be brief. Such a possibility may be based on the presence in system (2.3) of an energy localization effect [8], in which not all the energy of one of the mutually coupled systems is transferred to the other. It can be proposed that the possibility of existence of a similar quality in the system of Eq. (2.3) will depend on how much the transform $P(x)$ which changes $\Pi(x)$ to $\Pi^{*}(x)$ for the parameter range limited by the conditions $\gamma \geqslant 1, \mathrm{n} \geqslant 1$, differs from an identity.

We will note that for $\gamma$, $\sigma$ values for which $n=0$ such an approach leads to a different result. Thus we will consider a system with $\gamma=5.625, \sigma=7.5$ at $a=0.5, k=4, b=5, Z=$ 3. Such values of $\gamma$ and $\sigma$ insure the presence of only three sequences of equilibrium positions for system (2.3), $G_{o e}, G_{0 o}$ and $G_{1} i_{e}$

$$
G_{1 i_{e}}=\left\{x, y: y_{1}=y_{2}=0 ; x_{1}=\pi i_{e} / k, x_{2}= \pm x_{2}(\gamma, \sigma)\right\} .
$$

The equilibrium point sequences $G_{o o}$ and $G_{1} i_{e}$ are saddle points for the surface $\Pi=\Pi(x)$, while at the equilibrium positions $G_{o e}$ the function $\Pi(x)$ has a maximum, i.e., in accordance with Ernshaw's theorem all equilibrium positions are unstable. For the case considered the surface $\Pi=\Pi(x)$ cannot be reduced to the form of Eq. (2.12). The topological equivalent of Hamiltonian (2.4) in the tube $T(x, y)$ will be the Hamiltonian

$$
\begin{equation*}
I^{*}=\frac{1}{2}\left(y_{1}^{2}+y_{2}^{2}\right)+\frac{1}{2}\left(-x_{2}^{2}+\cos x_{1}\right) . \tag{2.16}
\end{equation*}
$$

Examination of the trajectories of Hamiltonian (2.16) reveals the absence of trajectories limited in the coordinates $x_{2}, y_{2}$. Thus, the limitation $n \geqslant 1$, where $n$ is determined from Eq. (2.7), issignificant and determines the possibility of representing Hamiltonian (2.4) in the equivalent (within tube $T$ ) form of Eq. (2.15).

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